

5-14-2012

## Small Scale Simulation Chamber for Space Environment Survivability Testing

Robert H. Johnson  
*Utah State University*

Lisa D. Montierth  
*Utah State University*

JR Dennison  
*Utah State University*

James S. Dyer  
*Space Dynamics Laboratory, Utah State University*

Alex Chanson  
*Utah State University & University of California, Berkeley*

Follow this and additional works at: [https://digitalcommons.usu.edu/mp\\_post](https://digitalcommons.usu.edu/mp_post)

 Part of the [Physics Commons](#)

---

### Recommended Citation

Johnson, Robert H.; Montierth, Lisa D.; Dennison, JR; Dyer, James S.; and Chanson, Alex, "Small Scale Simulation Chamber for Space Environment Survivability Testing" (2012). 12th Spacecraft Charging Technology Conference 14-18 May, 2012 Kitakyushu Japan. *Posters*. Paper 9.  
[https://digitalcommons.usu.edu/mp\\_post/9](https://digitalcommons.usu.edu/mp_post/9)

This Poster is brought to you for free and open access by the Materials Physics at DigitalCommons@USU. It has been accepted for inclusion in Posters by an authorized administrator of DigitalCommons@USU. For more information, please contact [digitalcommons@usu.edu](mailto:digitalcommons@usu.edu).







# Simulation Chamber for Space Environment Survivability Testing

Robert Johnson<sup>2</sup> and Lisa Montierth<sup>1</sup> Mentors: J R Dennison<sup>2</sup>, James Dyer<sup>1</sup>, and Ethan Lindstrom<sup>1</sup>

Utah State University, Logan, UT 84332-4414

<sup>1</sup>Thermal Vacuum Group Space Dynamics Laboratory

<sup>2</sup>Materials Physics Group Physics Department



## Space Environment Effects

The space environment can modify materials and cause detrimental effects to satellites. Some of these effects are change in reflectivity and emissivity, which lead to changes in thermal, optical, and charging properties. If these are severe enough the spacecraft will not operate as designed. The key to predicting and mitigating these deleterious effects is the ability to accurately simulate space environment effects through long-duration, well-characterized testing in an accelerated, versatile laboratory environment.

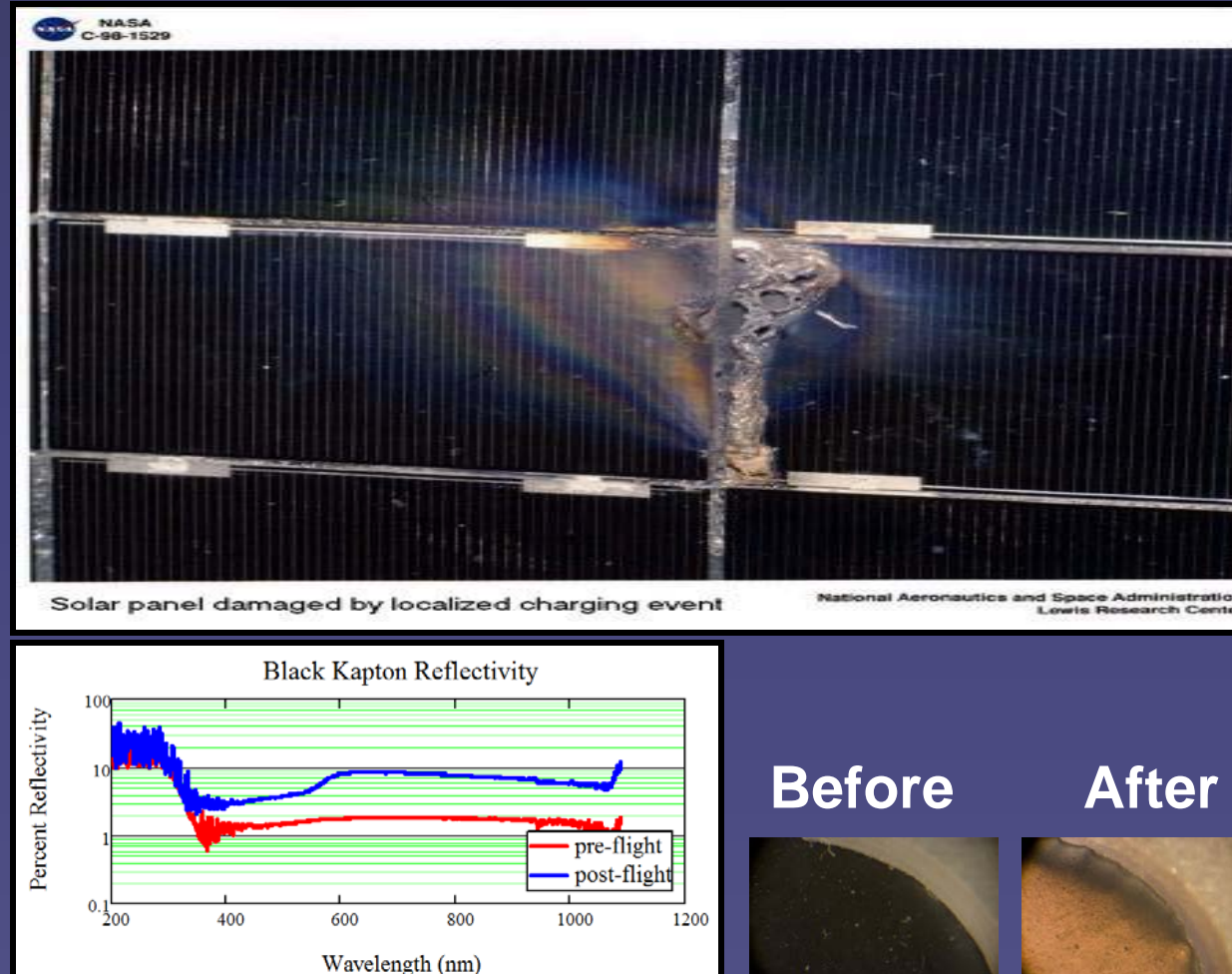


Fig. 6. (Top) Damaged Solar Panel due to overcharging. (Bottom) Photographs and UV/VIS/NIR spectra comparing pre- and post-flight samples from SUSPES II [Dennison].

## Abstract

A versatile vacuum system for long duration testing of materials modifications due to exposure to simulated space environment conditions has been designed and built. The chamber is particularly well suited for cost-effective tests of multiple small scale materials samples over prolonged exposure. Critical environmental components simulated include neutral gas [ultrahigh vacuum ( $10^{-9}$  Torr) to ambient], FUV/UV/VIS/NIR solar spectrum, electron plasma fluxes, and temperature. The UV/VIS/NIR solar spectrum is simulated using an external, normally incidence and collimated class AAA Solar Simulator source, with standard Air Mass Zero (AM0) filters to shape the incident radiation spectrum. This Xe arc discharge tube source has a 200 nm to 2000 nm range with up to four suns light intensity capability. Light intensity feedback is used to maintain the intensity temporal stability during the sample exposure cycle, with standard calibrated solar cells mounted internally on the sample mounting block. Incident FUV (far ultraviolet) intensity radiation is provided by Kr discharge line sources, with a primary emission line at 124 nm and secondary emission line at 117 nm with up to four suns intensity. This provides an adequate substitution for the solar FUV spectrum, which is dominated by the ultraviolet hydrogen Lyman  $\alpha$  emission line at 122 nm. An electron flood gun provides a uniform, monoenergetic ( $\sim 20$  eV to  $\sim 15$  keV) electron flux. Electron fluxes at the sample surface of  $<1 \mu\text{A}/\text{cm}^2$  to  $>1 \mu\text{A}/\text{cm}^2$  are continuously monitored during the sample exposure cycle, using a standard Faraday cup mounted on the sample block. The chamber maintains  $\leq 98\%$  uniformity of the electromagnetic and electron radiation exposure over a sample area of  $\sim 70 \text{ cm}^2$ . Samples are mounted on a rotatable OFHC Cu sample block with large thermal mass to minimize the differences in temperature between samples and thermal fluctuations during the sample exposure cycle. A controlled, uniform temperature range from 100 K to 450 K is achieved using a cryogenic reservoir and resistance heaters attached to the sample block. The sample carousel is attached to a standard rotational vacuum feedthrough, to allow  $355^\circ$  rotation of the samples relative to the incident fluxes. Reflectivity and emissivity are measured by extending a compact integrating sphere with a fiber optic connection to an external calibrated commercial UV/VIS/NIR spectrometer and an IR absorptivity/emissivity probe mounted on a linear translation stage toward the center of the chamber; each sample and *in situ* calibration standards are rotated under the probes in turn. An automated data acquisition system periodically monitors and records the environmental conditions, UV/VIS/NIR reflectivity, and IR emissivity of the samples *in situ* during the sample exposure cycle.

## Space Environment Characteristics

There are certain characteristics of the space environment that are critical for a true simulation. These critical characteristics are electron flux, electromagnetic radiation, vacuum, and temperature. The electron flux is critical because the solar winds through space bombard spacecraft. The electromagnetic radiation has many critical aspects in itself. As can be seen in figure 10, the sun has a very broad range covering from the Visual/Infrared to Ultra Violet, specifically the Hydrogen Lyman Alpha emission at 121.6 nm. A vacuum simulation is critical because space is a vacuum, meaning very few particles. The temperature is critical because it changes drastically depending on proximity to the sun. Things not covered by this chamber are photons/ions, and atomic oxygen.

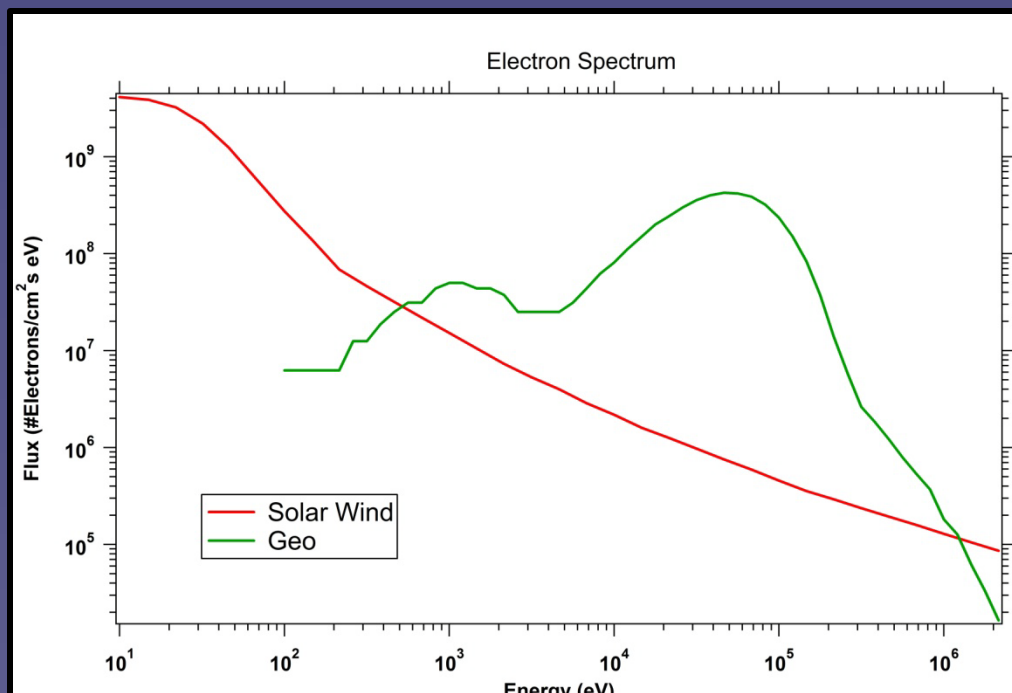
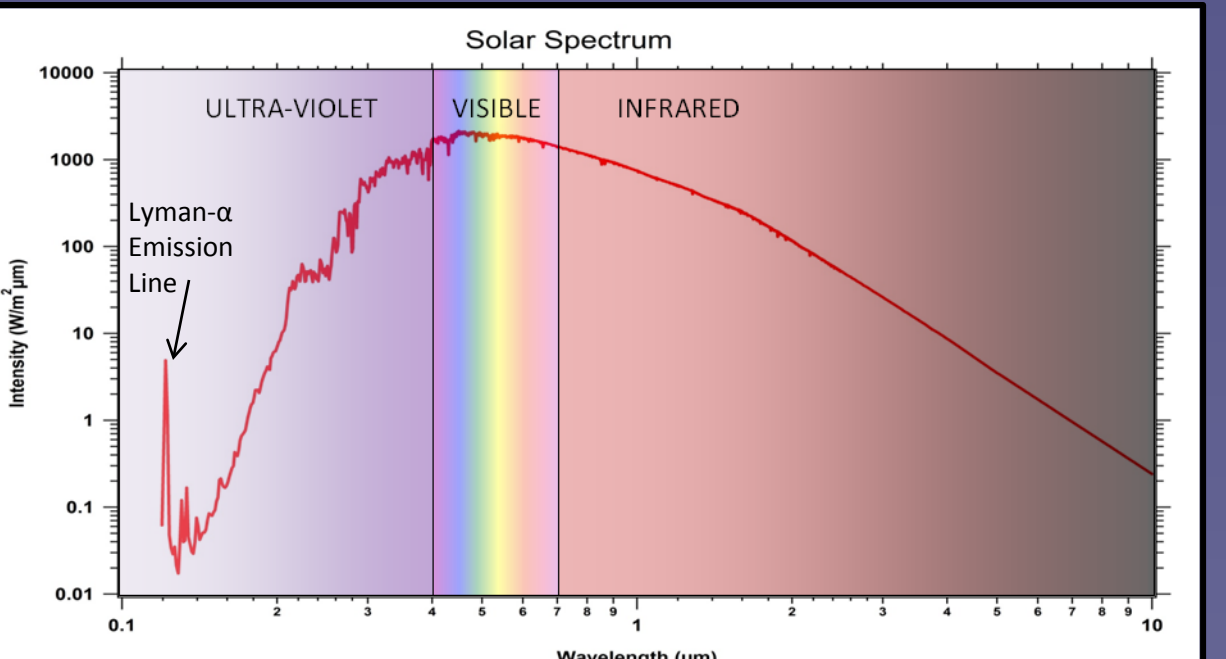


Fig. 8. (Left) Typical Space Electron Flux Spectra [Larsen].

Fig. 9. (Bottom Left) Solar wind and Earth's magneto-sphere structure.

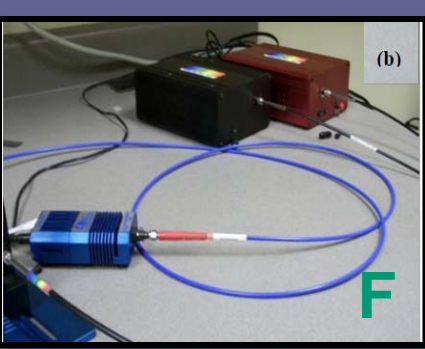
Fig. 10. (Bottom Right) Solar Electro-magnetic Spectrum.



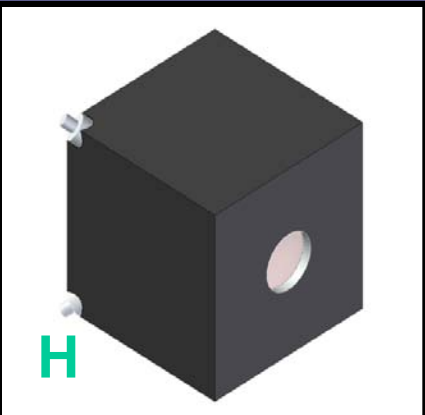
## In Situ Analysis Capability



**UV/VIS/NIR Reflectivity**—Two fiber optic spectrometers (F) measures reflectivity of UV/VIS/NIR (200-1080 nm) NIR (858-1700 nm) ranges with  $<1$  nm resolution.



**Integrating Sphere**—A 2.5 cm diameter integrating sphere (H) can be extended over the samples with a retractable probe linear translation stage (T). The sample stage can be rotated to position different samples under the probes. Light from a deuterium/W-halogen calibrated light source enters the integrating sphere through one fiber optic connection; reflected light from the sample exits through another fiber optic to spectrometers.



**IR Emissivity**—Measured with retractable probe (4  $\mu\text{m}$  to 15  $\mu\text{m}$ ) (G) mounted on probe translation stage.

**Calibration Samples**—*In situ* high and low reflectivity/emissivity calibration standards (N) are mounted behind the probe translation stage.



**Light Flux**—Continuously monitored with *in situ* photodiodes (I) on sample stage (M) equipped with filters to separately monitor NIR, VIS, UV intensities. Exterior sensor feedback used to regulate the solar simulator intensity.



**Electron Flux**—Continuously monitored with *in situ* Faraday cup (J) on sample stage (M).

**Temperature**—Monitored with platinum RTDs (K).

**Pressure**—Absolute pressure monitored with Convection and ion gauges (Y). Partial pressure measured with a Residual Gas Analyzer (Z).

## Experimental Test Chamber Design

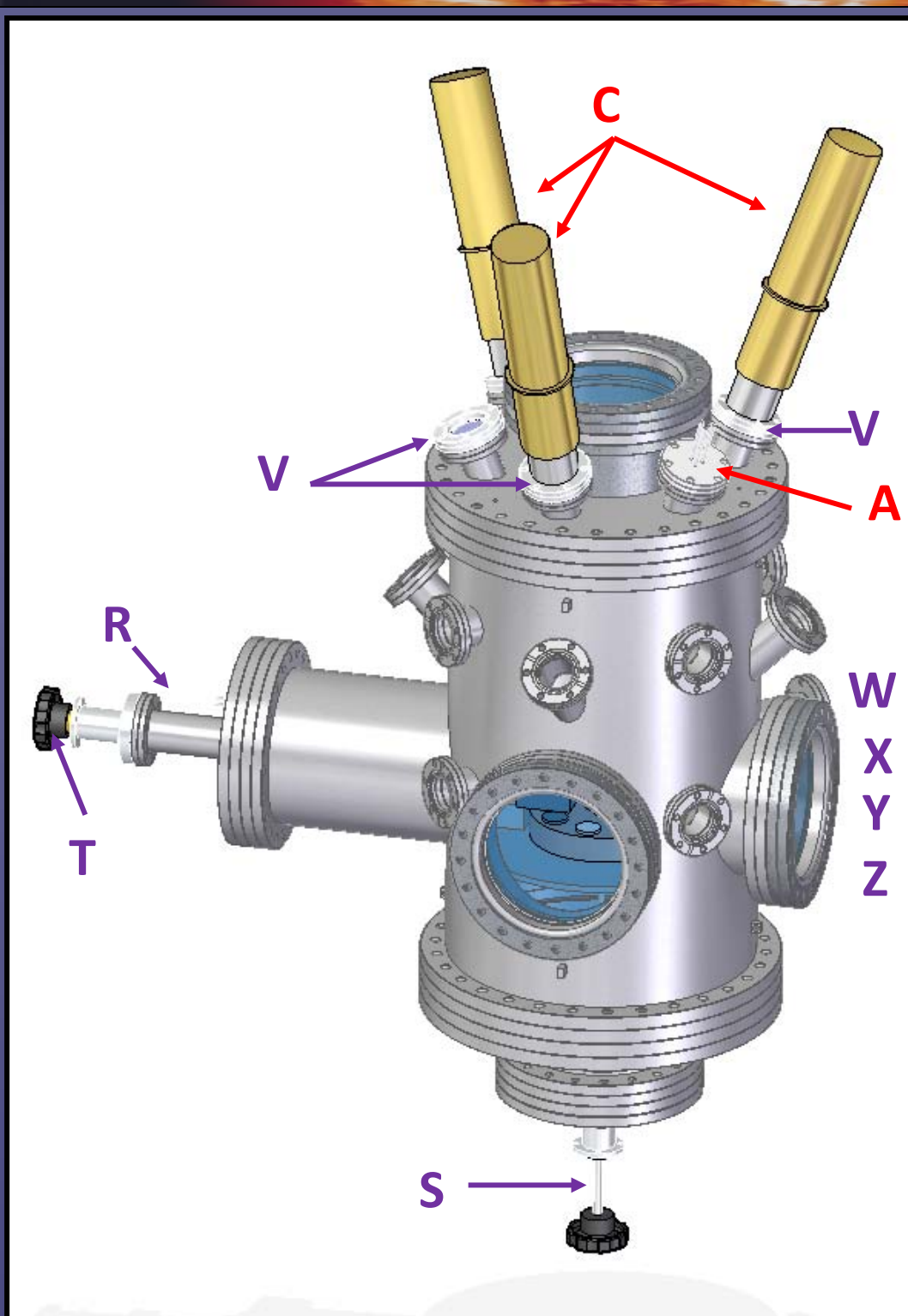


Fig. 1. Chamber Exterior View.

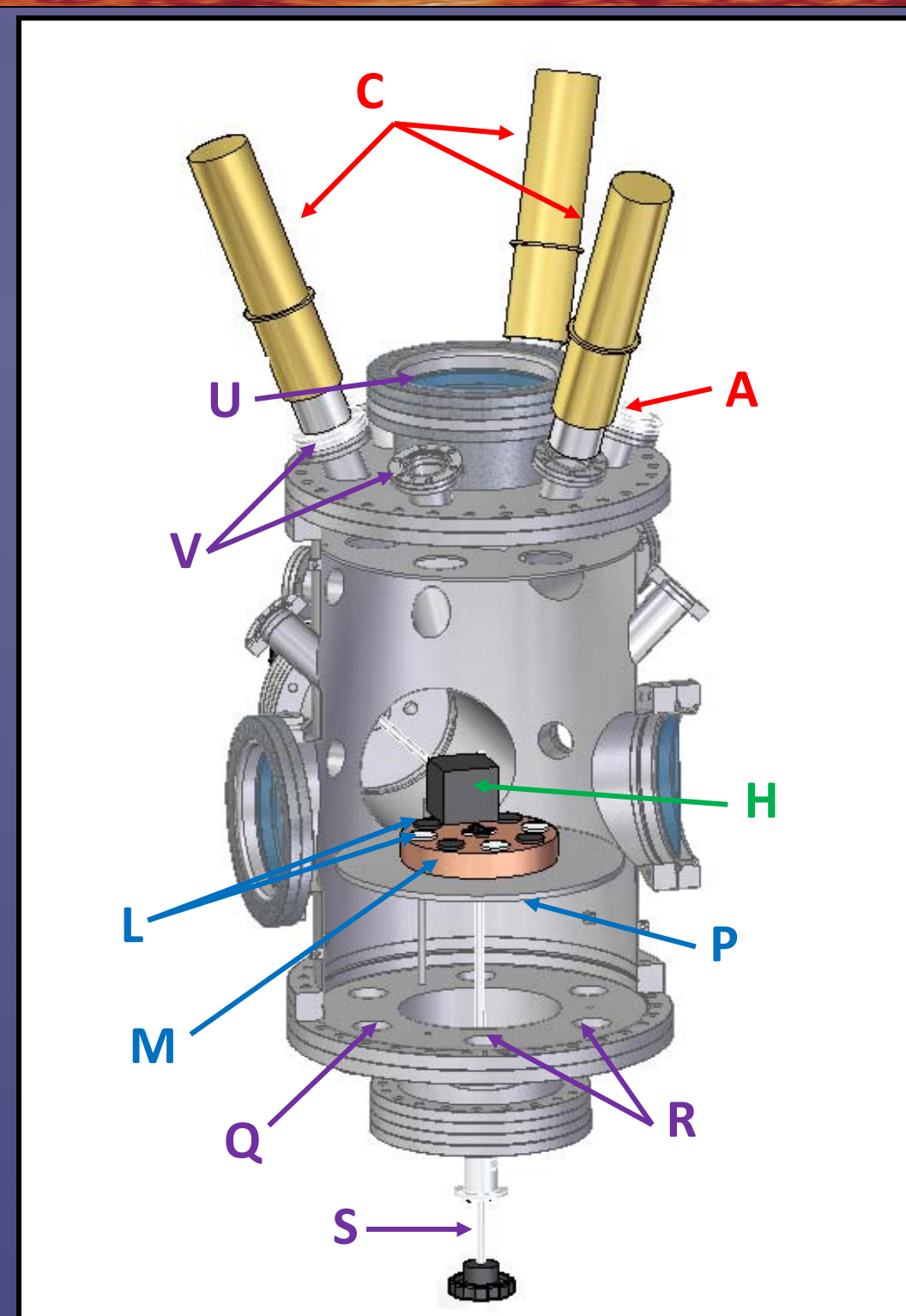


Fig. 2. Chamber Vertical Cutaway View.

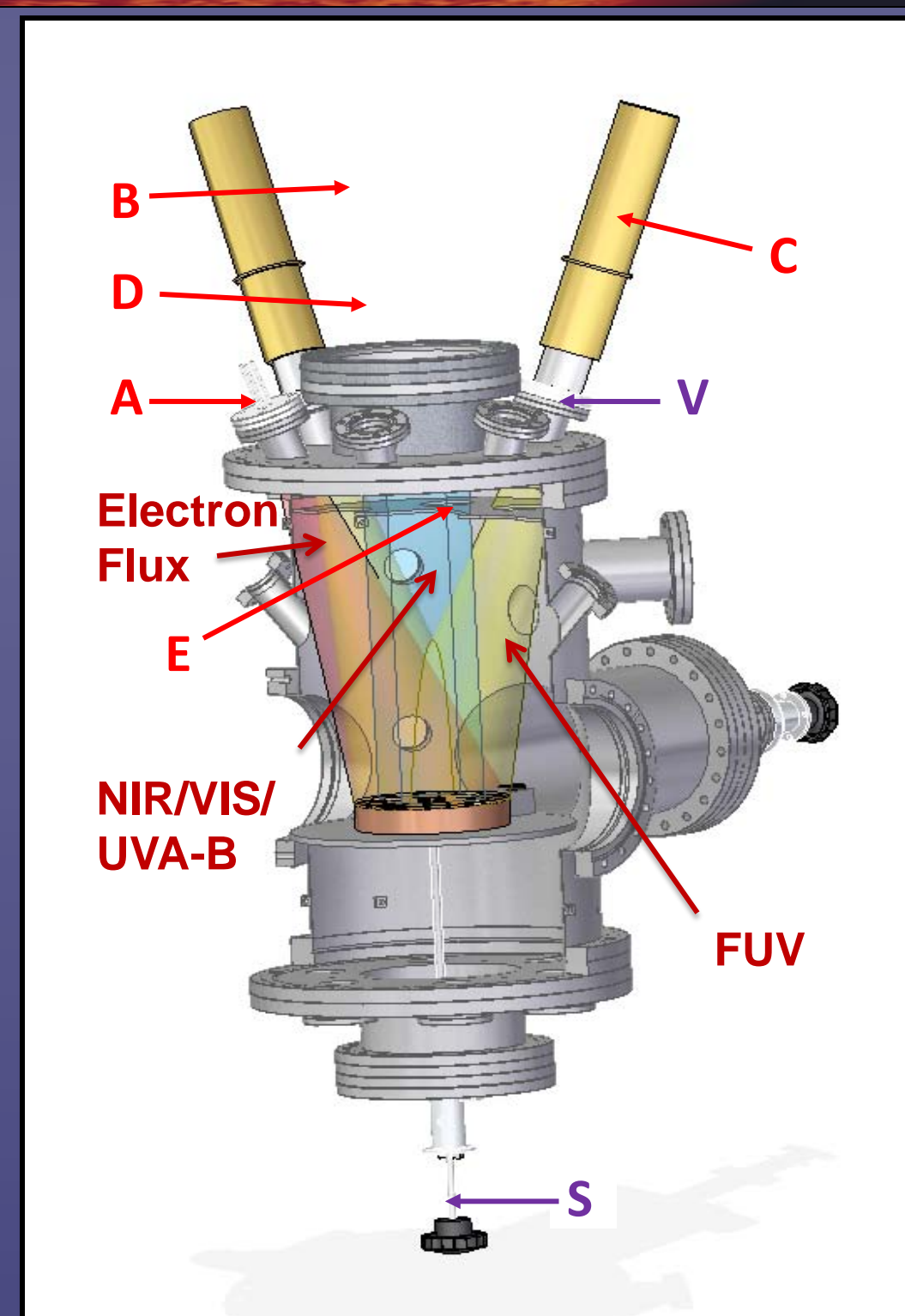


Fig. 3. Cutaway View of Beam Trajectories.

### Radiation Sources

- A Electron Gun
- B UV/VIS/NIR Solar Simulator
- C FUV Krypton Discharge Lamps
- D Air Mass Zero Filter Set
- E Flux Mask

### Analysis Components

- F UV/VIS/IR Reflectivity Spectrometers
- G IR Emissivity Probe
- H Integrating Sphere
- I Photodiodes—UV/VIS/IR Flux Monitor
- J Faraday Cup—Electron Flux Monitor
- K Platinum Resistance Temperature Probe

### Legend of Components

#### Sample Carousel

- L Samples
- M Rotating Sample Carousel
- N Reflectivity/Emissivity Calib. Standards
- O Resistance Heaters
- P Cryogen Reservoir

#### Instrumentation (Not Shown)

- Data Acquisition System
- Temperature Controller
- Electron Gun Controller
- UV/VIS/NIR Solar Simulator Controller
- FUV Kr Resonance Lamp Controller
- Spectrometers and Reflectivity Source

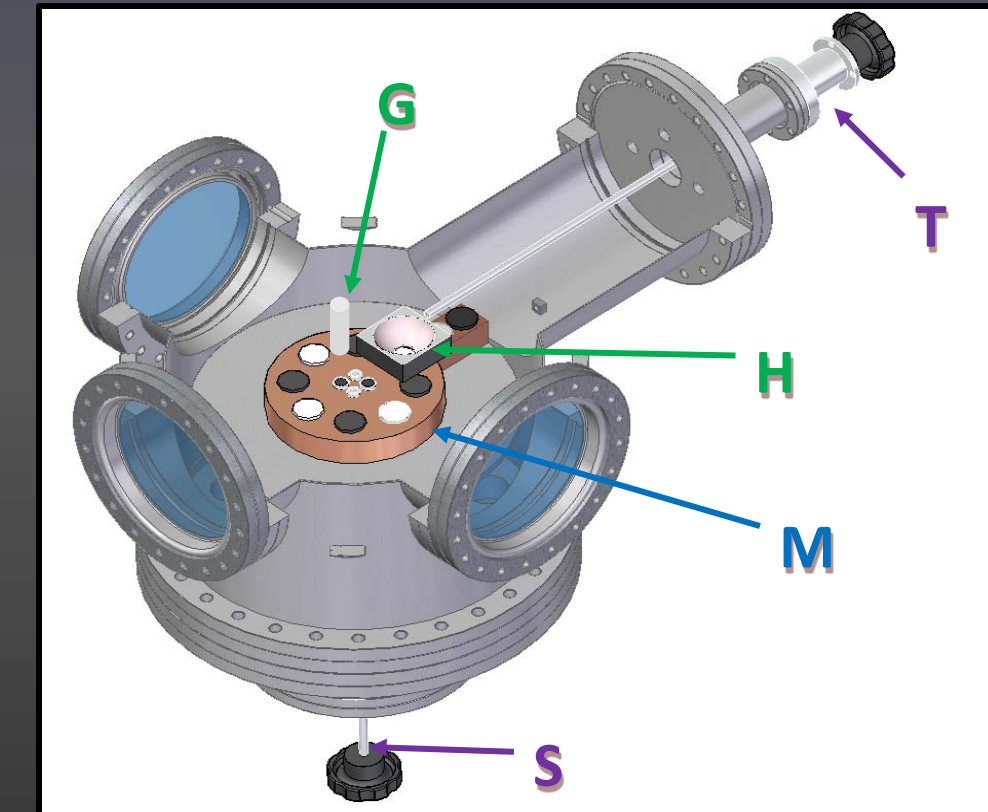
#### Chamber Components

- Q Cryogen Vacuum Feedthrough
- R Electrical Vacuum Feedthrough
- S Sample Rotational Vacuum Feedthrough
- T Probe Translational Vacuum Feedthrough
- U Sapphire UV/VIS Viewport
- V MgF UV Viewport
- W Turbomolecular/Mech. Vacuum Pump
- X Ion Vacuum Pump
- Y Ion/Convection Pressure Gauges
- Z Residual Gas Species



Fig. 4. (Right) Sample Level Cutaway View

Fig. 5. (Left) The chamber can be reconfigured as a radiation source for other test chambers by removing the sample stage flange and bolting the upper source components to other SDL/USU chambers.



## Space Simulation Capabilities

Versatile ultrahigh vacuum test chamber provides controlled temperature and vacuum environment with stable, uniform, long-duration electron and UV/VIS/NIR fluxes at up to 4 times sun equivalent intensities for accelerated testing for a sample area of 8 cm by 8 cm. Particularly well suited for cost-effective tests of multiple small scale materials samples over prolonged exposure.

**Electron Flux**—Electron flood gun (A) provides  $\leq 5 \cdot 10^6$  electrons/ $\text{cm}^2$  ( $\sim 1 \mu\text{A}/\text{cm}^2$  to  $1 \mu\text{A}/\text{cm}^2$ ) flux needed to simulate the solar wind at more than the 100X cumulative electron flux. Mono-energetic energy range is  $\sim 0.05$  to  $15.00 \pm 0.01$  keV. Gun provides a  $>98\%$  uniform flux distribution over the full sample area, with "hot swappable" filaments for continuous exposure over the entire long duration testing. The electron gun was custom designed at USU after work by Swaminathan [2004].

**Infrared/Visible/Ultraviolet Flux**—A commercial Class AAA solar simulator (B) provides NIR/VIS/UVA/UVB electromagnetic radiation (from 200 nm to 1700 nm) at up to 4 times sun equivalent intensity for accelerated testing over an area of 80mmX80mm. Source uses a Xe discharge tube, parabolic reflector, collimating lens, and standard Air Mass Zero filters (D) to match the incident radiation spectrum to the solar spectrum. Xe bulbs have  $>1$  month lifetimes for long duration studies.

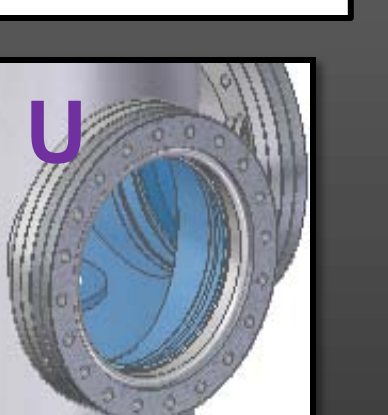
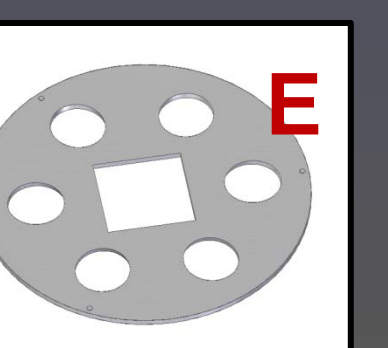
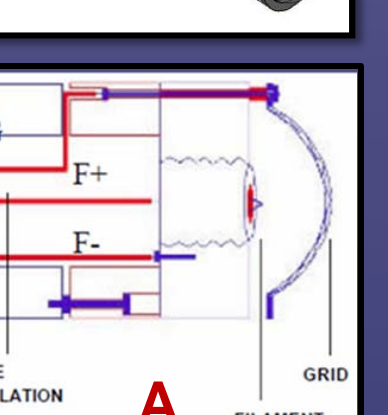
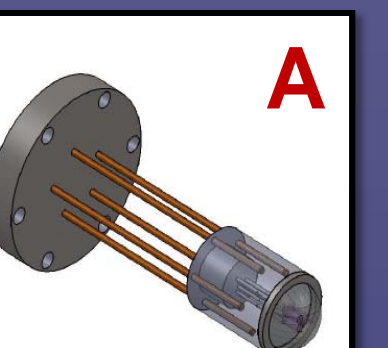
**Far Ultraviolet Flux**—The Kr resonance lamps (C) provide FUV radiation flux (ranging from 10 to 200 nm) at 4 times sun equivalent intensity. Three lamps oriented  $120^\circ$  apart provide  $>98\%$  flux uniformity. Lamp's emission lines reproduce the H Lyman- $\alpha$  line (121.6 nm) that dominates the solar FUV spectrum. Kr bulbs have  $\sim 3$  month lifetimes for long duration studies.

**Flux Mask**—Flux mask (E) located near the chamber's top ports restricts the flux boundaries to the sample stage, limiting equipment exposure and reducing scattering to accommodate uniform exposure. Can be readily modified for different sample geometries.

**View Ports**—Solar simulator UV/VIS/NIR light passes through sapphire viewport (U). Krypton source FUV light passes through a Magnesium Fluoride window (V). Additional viewports allow visual inspection.

**Vacuum**—Chamber uses standard mechanical and turbomolecular pumps (X) for roughing and an ion pump (Y) for continuous maintenance-free operation (base pressure  $<10^{-5}$  Pa).

**Temperature**—A temperature range from 100 K to 450 K is maintained to  $\pm 2$  K by a standard PID temperature controller, using a cryogenic reservoir (Q) and resistance heaters (P) attached to a large thermal mass sample stage (M).



## Versatile Sample Holder Design

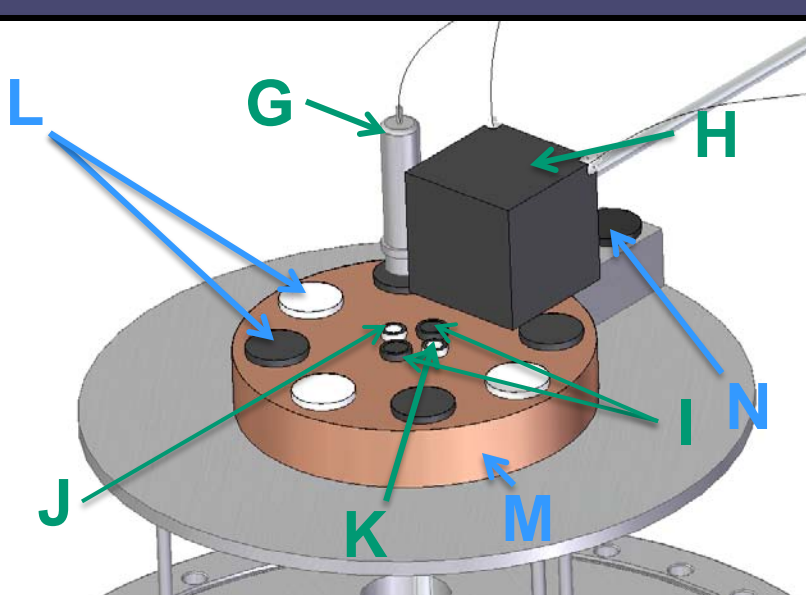


Fig. 7. View of Sample Carousel and Probe Translation Stage.

**Sample Stage**—Sample stage (M) connected to  $355^\circ$  rotary feedthrough (S) to position samples under probe translation stage (T) and enhance flux uniformity by periodic rotation. Sample stage shown has six 2.5 cm diameter samples (L) plus flux sensors (I, J, K); alternate configurations have up to one 10 cm diameter sample. Uniform temperature over  $\sim 100$  K to 450 K controlled using attached cryogenic reservoir (P) and resistance heaters (O). Large thermal mass helps maintain stable thermal.

## Acknowledgements/References

- Larson, D.E et al., *Solar Wind 8 Conf. Proc.*, submitted (1996a).
- Dennison, J R, "Charge-Enhanced Contamination and Environmental Degradation of MISSE-6 SUSPES Materials," IEEE Transactions on Plasma Science, February 2012, Vol. 40
- Dennison, J R and Dyer, James S., "Proposal to Perform Survivability Testing of Radiator Coatings for Solar Probe Plus", May 27, 2011.
- Swaminathan, Prasanna V., "Measurements of Charge Storage Decay Time and Resistivity of Spacecraft Insulators," Masters Thesis, Utah State University, Logan, UT 2004.